0: ½ MPC ~ 1.5 Mly, the Milky Way,

including its two *Magellanic Cloud* irregular spirals and many spheroidal dwarf galaxies and GCs.

Most galaxies as seen in small telescopes from a suburban backyard will show up as just faint patches or stellar points surrounded by tiny halos; What interests me though, is not so much what I'm able to catch of details in the individual objects, but rather the large-scale location and distribution of the galaxies, that is: how the groups can be seen "wide field" on the celestial dome, as I zoom out from our Milky Way...

Below I've listed the Messier galaxy objects, arranged according to their distance from our own Milky Way galaxy:

1: 2 Mpc - 5 Mly, the Local Group,

The Milky Way – Andromeda – Triangulum galaxies with their satellites, plus the *Sculptor Group* (N55 at ~10 Mly).

2: 6 Mpc ~ 20 Mly, the Local Supercluster, The M81 - CVn – M83 – CenA groups, plus the Leo Groups (several Messiers at ~40 Mly).

- 3: 20 Mpc \sim 70 Mly, the Virgo Supercluster, The UMa-Coma-Virgo groups, all in the background at \sim 50 Mly.
- 4: 90 Mpc ~ 300 Mly, the Surrounding voids and supercluster filaments,

for instance the Taurus Void and Perseus-Pisces wall.

5: And beyond...

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M NGC LOCAL GROUP (5 Mly) Autumn
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M31 N224 S And Great Andromeda M32 N221 E And Dwarf to M31

M110 N205 E And Dwarf to M31

5: Galaxies [40]

M33 N598 S Tri Triangulum/Pinwheel

M NGC LOCAL SUPERCLUSTER (Near~ 15 Mly)

M81 N3031 SA(s)ab UMA Bode's M82 N3034 IO UMa Cigar

M94 N4736 (R)SA(r)ab CVn CrocEye Canes-I

M64 N4826 (R)SA(rs)ab Com Black Eye

M83 N5236 SAB(s)c Hya S.Pinwheel Cen-A

M NGC LOCAL SUPERCLUSTER (Far~ 30 Mly)

M101-Cloud

M101 N5457 SAB(rs)cd UMa Pinwheel
M51 N5194 SAbc CVn, Whirlpool
M63 N5055 SAbc CVn Sunflower

 M66
 N3627 SAB(s)b Leo
 E Triplet

 M65
 N3623 SAB(rs)a Leo
 E Triplet

 M96
 N3368 SAB(rs)ab Leo
 W Triplet

 M95
 N3351 SB(r)b Leo
 W Triplet

M105 N379 E1 Leo W Triplet

M NGC VIRGO SUPERCLUSTER (60-70 Mly)

- VIRGO-I Core
- UMA-I Association (SUPERGROUP)
- COMA-I Association (SUPERGROUP)

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5: Galaxies 40 -- continued]
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M NGC VIRGO-I CORE (60 Mly)
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M87 N4486 cD pec
                  Virgo A
                             Wall W
M86 N4460 E2 Vir
                  Markarian Wall W
                  Markarian Wall W
M84 N4374 E1 Vir
M58 N4579 SAB(rs)b Vir
                             Wall E
                             Wall E
M59 N4621 E5 Vir
                             Wall E
M60 N4649 E2 Vir
                   Virgo C
M49 N4472 E2 in Vir Virgo B Outskirts
M61 N4303 SAB(rs)bc Virgo S Outskirts
M85 N4382 SA0+(s)pec Com
                           Outskirts
M88 N4501 SA(rs)b
                    Com
                             Hook N
M91 N4589 SB(rs)b
                    Com
                             Hook N
M89 N4552 E0-1
                    Vir
                             Hook S
M90 N4569 SAB(rs)ab
                    Vir
                             Hook S
M98 N4192 SAB(s)ab
                           Triangle
                    Com
M99 N4254 SSA(s)c
                    Virgo N Triangle
M100 N4321 SAB(s)bc
                            Triangle
                     Com
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M NGC COMA-I Association (60 Mly)

-- N4725,N4565 --<no Messier galaxies>

M NGC UMa-I Association (70 Mly)

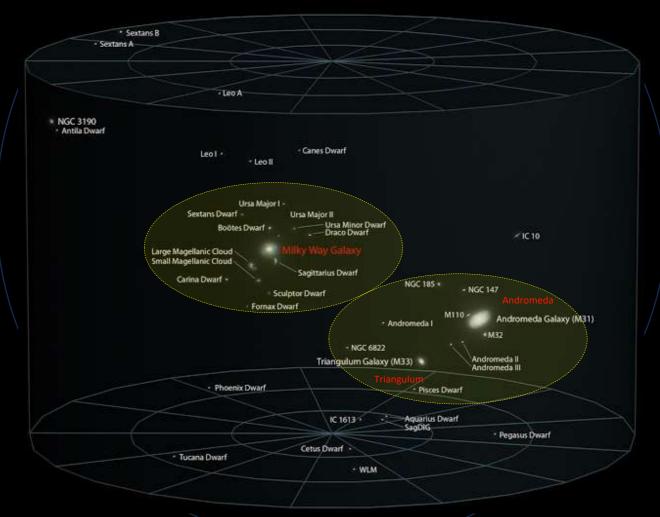
M106 N4258 SAB(s)bc CVn M108 N3556 SAB(s)cd UMa M109 N3992 SB(rs)bc Uma

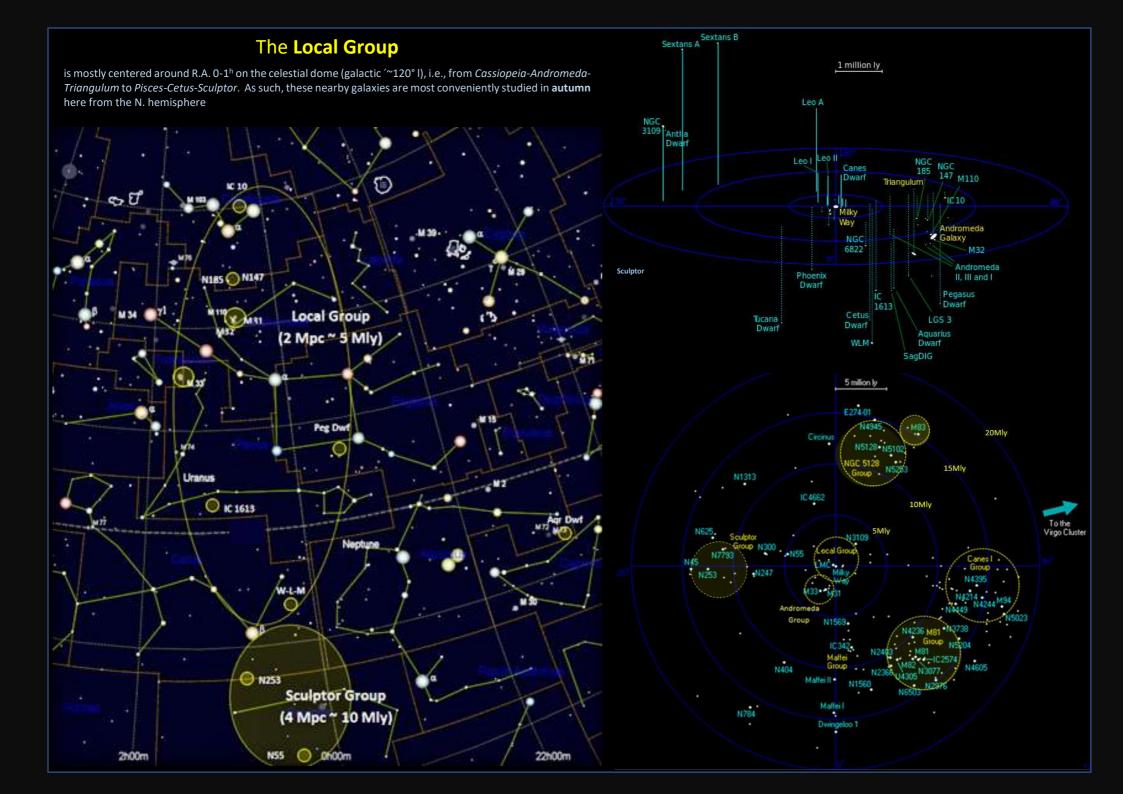
```
M74 N628 SA(s)c Psc Phantom 30 Mly
M104 N4594 SA(s)a Vir Sombrero 30 Mly
M102 N5866 SAO Draco Spindle 50 Mly
M77 N1068 (R)SA(rc)b Cetus 50 Mly
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Step 1: 2 Mpc ~ 5 Mly The **Local Group**

The Milky Way – Andromeda – Triangulum galaxies with their satellites,

plus, the *Sculptor Group* (N55 at ~10 Mly).

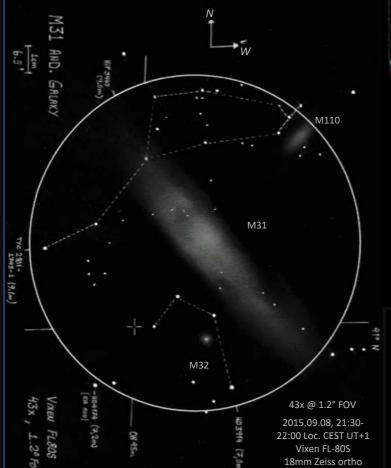


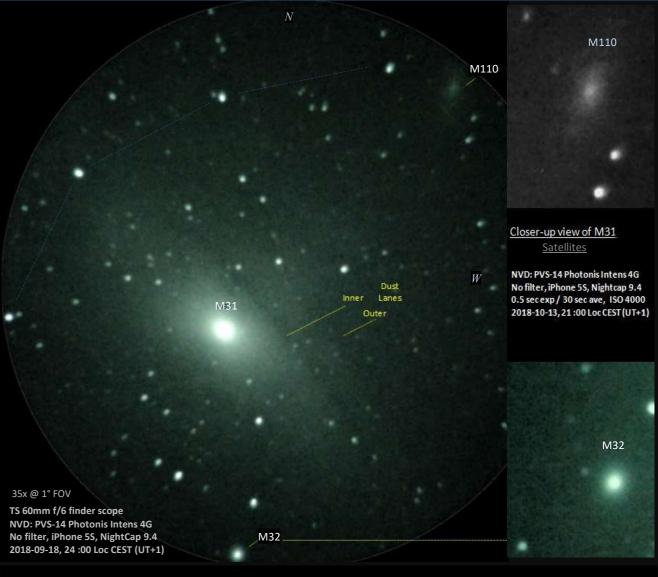


Messier *M31 – M32 – M110*

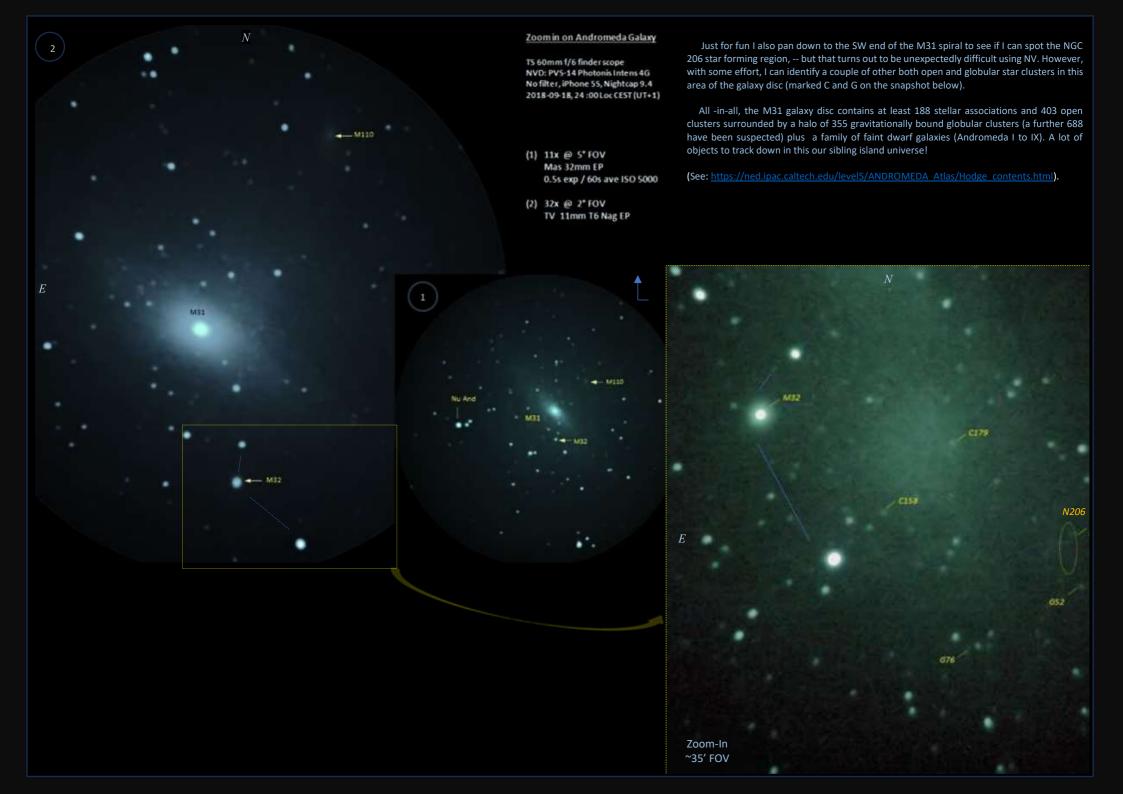
Here's a drawing I made back in 2015 (Vixen FL-80S, 18mm ortho, 43x mag.; see: https://www.cloudynights.com/topic/523941-classic-messier/?p=7085247). As seen from our own galaxy, M31 is inclined 77° with the nearest edge of the disc upwards towards the NW, and a clockwise rotation of the spiral (as seen "bottom-up") with the trailing arms receding at the NE tip and approaching at the SW end. The core of M31 shows a graduation from a bright circular nucleus, over a surrounding elliptic disc that progressively fades into an elongated faint halo. The galaxy seems brighter and more extended towards the SE, while the NW edge is more sharply delineated due to the dark dust lanes found here.

In my 3" scope the M32 satellite looks like a round "wooly" 8^m star S of the core of M31, while M110 shows up as a fainter but larger and elliptical haze towards the NW. M32 has a central supermassive black hole and is probably the compact core of a disc galaxy that has been torn apart by M31, while M110 is a more typical dwarf elliptical in a further out orbit around M31, though still under influence by the "mothership" as evidenced by recent star forming activity.





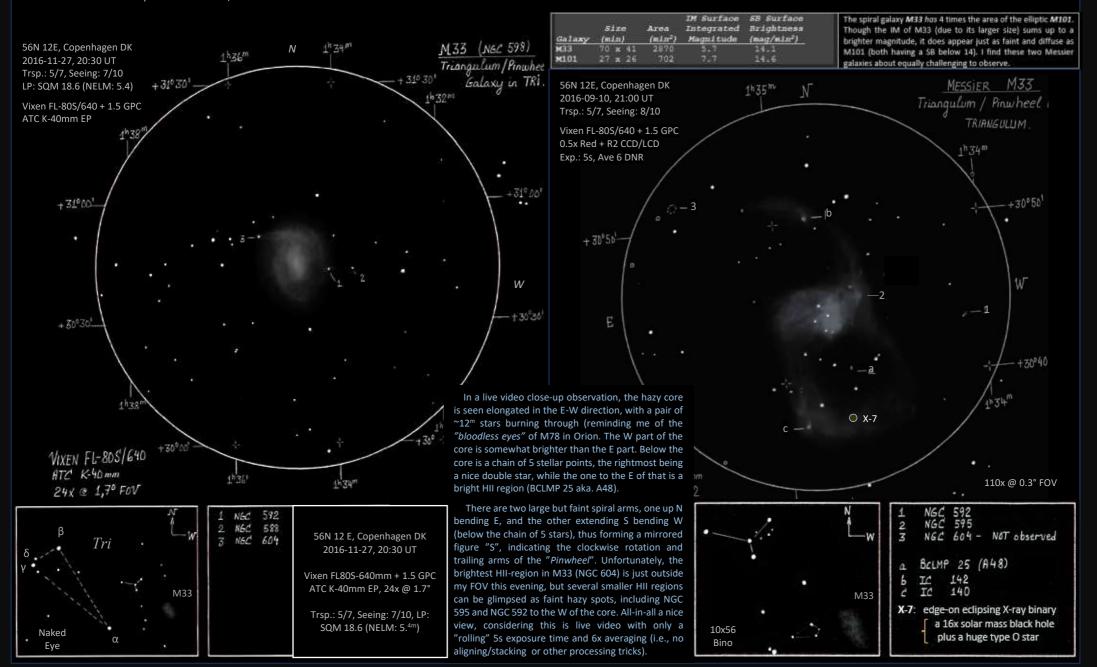
Now, three year later, I'm out again on a late evening in August (2018-09-18, 24:00 Loc CEST, UT+1) to observe some HII-regions in Perseus with my Night Vision monocular; Before calling it a night, I decide to do a fast zoom-in on the Andromeda galaxy, which is now well placed at 64° Alt. in the SE, and closing in on the meridian. Using 11x @ 5° field I can trace the spiral across ~2° of the night sky, and zooming in to 32x @ 2° field I can clearly spot the two closest satellites (M32, M110) as well as the inner dark dust lane in the spiral disc. Finally zooming in to 35x @ 1° field, both inner and outer dust lanes in M31 are now evident, and I can get a better close-up look at the M32 and M110 satellites.



M33 (NGC 598), the Triangulum (or "Pinwheel") galaxy

It's mid-September (2015), the hour around midnight. A calm night after rain, but still with drifting high clouds, so the transparency is varying, -- from good in the cloud holes to damn lousy when a cloud drifts by... The seeing is steady though, so I give M33 a try with my Zeiss 10x56 bino (FOV 6°). I have mounted the bino using a Manfrotto clamp on a Zeiss dovetail rail that fits my motorized CZJ TM/2V mount. This leaves me with both hands free to draw.

I first pan the bino from *Alpha TRI* to M33. The galaxy is now clearly visible in the FOV of the bino, filling the whole E side of an equilateral triangle defined by three 8^m stars (H9483, HD9269, TYC2293-912-1). It shows up as a faint diffuse grey haze, elongated in the NNE – SSW direction, with an ever so slightly brighter central area, but with no well-defined outline or internal structure. It reminds me of the naked eye view of *M31 AND*, on a dark, clear night. It is interesting to compare M33 TRI with another relatively faint Messier galaxy in Andromeda: M101:



M33 (NGC 598), the Triangulum / "Pinwheel" galaxy

is in orbit around the *Great Andromeda Galaxy* M31. Currently it is at a distance of almost 3 Mly from the MW, -- a little further out than M31 itself - but possibly linked to "the mothership" by a broad stream of HI gas and faint stars (see: https://iopscience.iop.org/article/10.1088/0004-637X/763/1/4).

As seen from the MW, M33 is a nearly face on clockwise rotating spiral ca 30% the size of M31. The nucleus of M33 is without a bar or central bulge, but the galaxy features two major trailing spiral arms: the northern arm containing four large HII regions (NGC 604,588,592,595) plus 10 smaller nebulous IC objects, while the southern arm exhibits ~143 large star clouds and OB-associations of young, hot stars (see: http://astronomy-mall.com/Adventures.In.Deep.Space/M33.HII-Star.Clouds.html).

Here's a "zoom-in" on M33 using NV (2010-11-02, suburban NELM 5.8, 92% humidity with drifting low cumulus). Night vision is obviously not the best choice with respect to resolution of galaxies (I prefer glass and/or EAA live video), but NV can help to at least identify these objects in high LP conditions:

The brightest HII region in M33 is NGC 604 at the tip of the N spiral arm, which rivals the *Tarantula Nebula* in our own *LMC*, as well in size as in star-forming activity; NGC 604 is easily spotted in even small amateur telescopes, as can be seen in my drawings and snapshot here. M33 harbors no central supermassive black hole, but it does in its S spiral arm have an extraordinary massive X-ray eclipsing binary, composed of a stellar black hole + a luminous type O-I blue supergiant (position marked on my drawing: X-7). M33 is surrounded by ~54 globular clusters (with 122 more currently suspected); C39 is the brightest GC in M33, but still very faint at ~16^m as seen here from our MW.

